

# Physics Division Overview

Jim Siegrist

Director's Annual Review November 8-9, 2005

# A Center of Excellence for the HEP Community



- Physics Ideas from LBNL
  - Asymmetric B Factory to probe CP violation in quark sector
  - Supernovae to measure the acceleration of the Universe
- Innovation at LBNL
  - Time Projection Chamber (TPC)
  - ASICs for silicon vertex detectors and pixel detectors
  - CCDs for space and ground-based astronomy
  - Analog Transient Waveform Digtiizer for non-Accelerator experiments
- Instrumentation developed at LBNL
  - BaBar: Silicon Vertex Tracker, Cerenkov Ring Detector, Trigger
  - CDF: Central Outer Tracker, Run II silicon
  - ATLAS: silicon strip modules and pixel detectors

Creativity, Ingenuity & Technical Capability

## Infrastructure is Highly Leveraged



- Outstanding faculty supported by UC Berkeley
- Small but dedicated full-time scientific staff
- Excellent technical resources
  - Computing Division (NERSC)
  - Engineering Division (e.g. IC design)
  - Large machine shops, clean room facilities
- Direct support from the lab via LDRD
- Synergy with Nuclear Science and Accelerator Divisions

# LBNL Contributions Enhance University Collaborations



- LBNL collaborates closely with the University community:
  - ✓ Shared equipment and infrastructure for chip design, silicon detector systems
  - ✓ Engineering expertise in advanced electronics, instrumentation and mechanical design
  - ✓Integration of theory with experiment
  - ✓ Computing expertise and operations support from NERSC

## **Program Overview**



**Accelerator Experiments** 

□ Present: BaBar, CDF

□Imminent: ATLAS

□R&D: ILC

■ Non-Accelerator Experiments

□ Present: KamLAND, SCP/SNF

□R&D: SNAP, APEX-SZ, South Pole Telescope

□Incubating:  $\theta_{13}$  at a reactor, CMB polarization

**□** Community Service

**□PDG**, Quarknet, Leadership

SNAP (Dark Energy) and ATLAS (EWSB) are our highest priorities

#### In this review



- You will see how
  - New theoretical ideas from LBNL challenge current and future experiments
  - —Current experiments at the Tevatron and later ATLAS at the LHC will probe the energy frontier and the limits of the standard model
  - —Our programs in supernova cosmology and the cosmic microwave background are exploring the 95% of the universe that lies beyond the standard model
  - A reactor-based experiment can initiate the next phase of neutrino physics
  - —The Particle Data Group provides an indispensable service as the repository and evaluator of results in high energy physics
  - —New ideas in instrumentation can form the technical basis for ILC detectors

# **Present and Future Program**



#### Present Program

- CDF is studying top cross section and mass
- Measurements of CKM parameters & B<sub>s</sub> mixing at BaBar & CDF
- New SNe found in intermediate redshift surveys
- Supernova Factory entering operation
- First observation of Geo neutrinos in KamLAND; 4 π arm completed
- APEX-SZ nearing first light

#### <u>Future Program – Centerpieces</u>

- ATLAS pixel final assembly underway
- Berkeley role as lead of west coast ATLAS analysis center being developed
- Substantial progress on SNAP/JDEM sensor R&D

#### <u>Future Program – In Development</u>

- Instrumentation R&D, detector concept development, and Physics studies for ILC
- R&D towards a new reactor experiment at Daya Bay

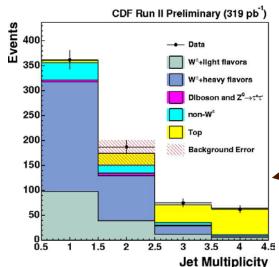


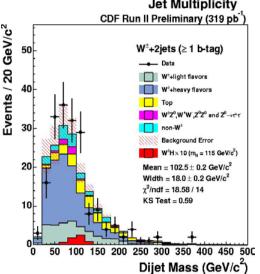
# Present Program

# LBNL Role in CDF Physics Program



#### **Electroweak + Top Physics**



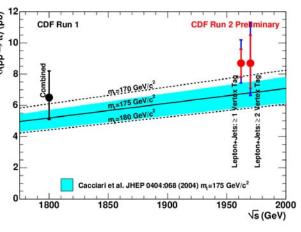


Top cross section
H. Bachacou (PHD thesis)

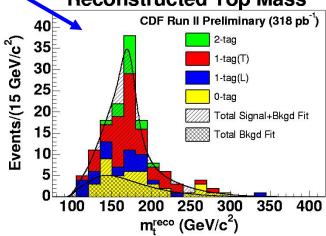
Top signal in I + □ 3 jets

Top Mass measurement E. Brubaker (PHD thesis) with others

Higgs Search
pp → W H
with H→ bb
limit on σ





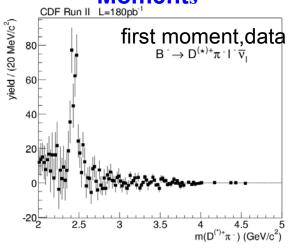


$$M_{top} = 172.5 \pm 3.9 \text{ GeV}$$

# LBNL B Physics Program





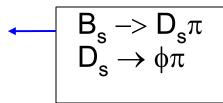


Measuring CKM parameters

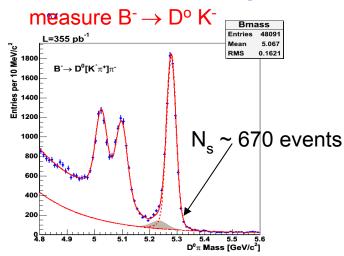
# CDF Fall B<sub>s</sub> Mixing

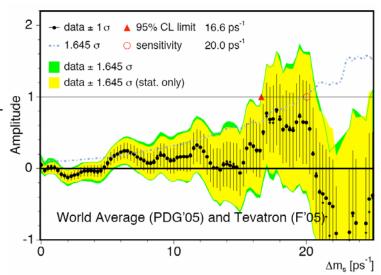
Limit: 7.9 ->8.6 ps<sup>-1</sup>

Sensi.: 8.4 ->13.0 ps

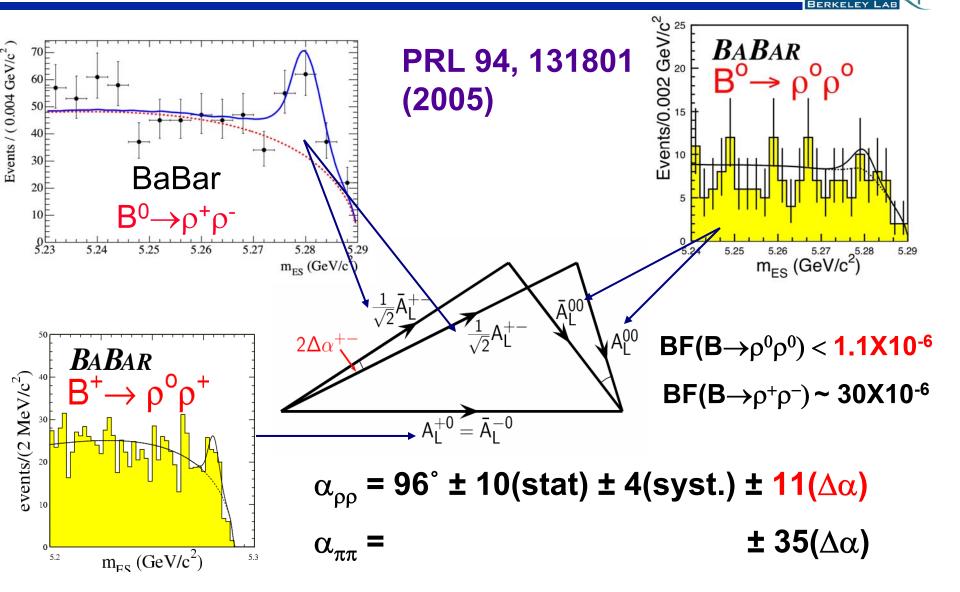


#### **Prospects for angle**





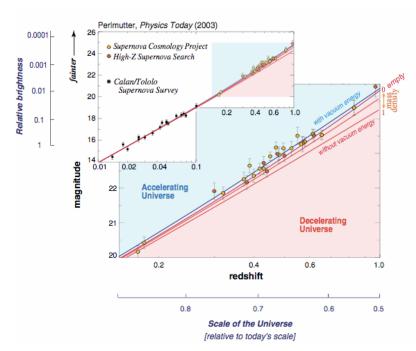
# VV pioneers: extracting Δα from B→ρρ (Gritsan, Groysman, Mir)



# Pioneers in Supernova Cosmology

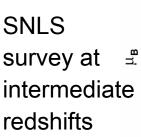


#### Supernova Cosmology Project

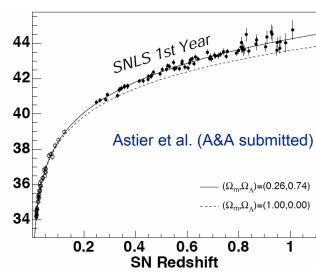


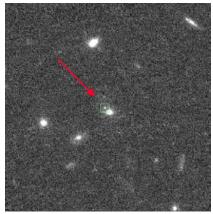
LBNL discovery of the acceleration of the universe established the new field of supernova cosmology and, more generally, dark energy studies

#### **Current Efforts**



Major HST SN search in high redshift clusters (z ≥ 1)

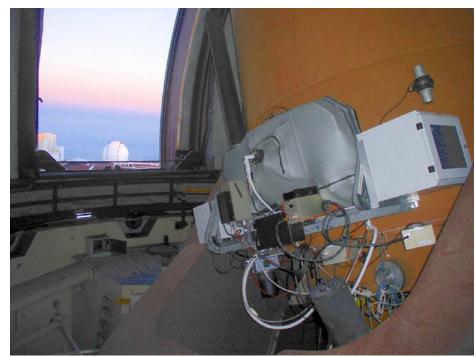




SN la discovered in Aug with HST/ACS in galaxy cluster at z=1.02

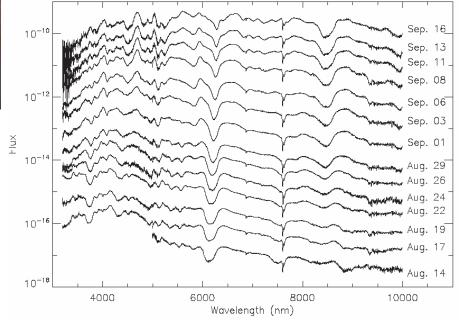
# SuperNova Factory

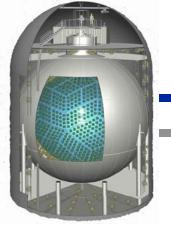




SNIFS spectral time series of the Type Ia supernova SN 2004dt (with arbitrary offset; no absolute calibration)

Remotely operated spectrograph (SNIFS) on the University of Hawaii 2.2m telescope (w/ Keck and Subaru in background). To date it has netted 65 SNfactory SNe plus as many community SNe.





#### KamLAND in 2005

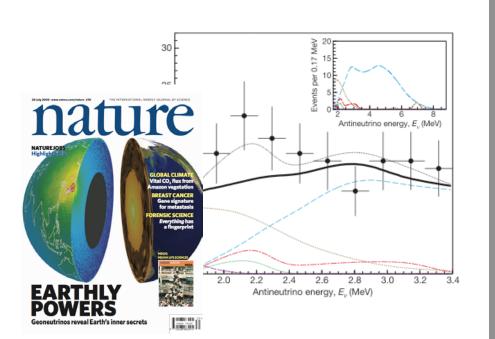


#### **Continuing Discoveries**

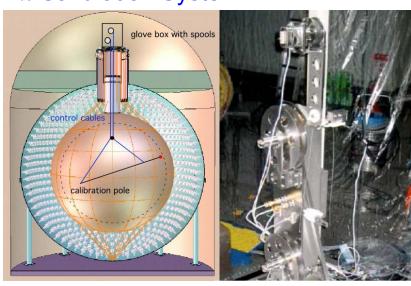
2003 Observation of  $\overline{\nu}_e$  disappearance

2004 Evidence for  $\overline{v}_e$  oscillation

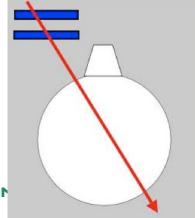
**2005** First Observation of Geoneutrinos



# Completed Construction & Testing of $4\pi$ Calibration System



#### **Muon Tracker Construction**





AWRENCE BERKELEY NATION

## **CMB** and Cosmology



- Integrated program combines effort at LBNL + UCB
- Strong program in theory, data analysis, algorithms
  - MAXIPOL, Planck
  - Collaboration with NERSC
- New instrumentation enables new experiments
  - APEX-SZ, South Pole Telescope
    - Galaxy Cluster Search probe Dark Energy
  - POLARBEAR design ripe for construction
    - CMB Polarization probe Energy Scale of Inflation
- LBNL leads readout development
- Significant funding through campus for joint program

# Atacama Pathfinder Experiment (APEX-SZ)





- 16,500 feet in Chilean Andes.
- 12m on-axis ALMA prototype

#### Berkeley SZ Receiver:

- 330 Bolometer array
- Discover 4000 Clusters/2yrs
  - Mass limit >  $4x10^{14} M_0$
- First Light Spring 2006
- LBNL responsible for readout

#### UC Berkeley/LBNL, MPI-Bonn/Munich, Cardiff

Galaxy Cluster Search - probe Dark Energy

#### Future:

- South Pole Telescope readout
- POLARBEAR (B mode polarization)

# Theory Group:increasing focus on LHC

Hall et al. – "improved naturalness" with nonstandard electroweak sectors implies dramatic signals @LHC:

- Mirror World related to SM quanta by discrete symmetry OR
- Two Higgs doublet model with strongly coupled heavy Higgs

Nomura et al. – reconciling MSSM with LEP lower limit in  $m_H$  implies  $\sigma(LSP - Nucleon) > 10^{-44} cm^2$  ensures observable LSP - dark matter signal @LHC

Bauer & Schwartz in progress – apply B-physics tool, SCET = Soft Collinear Effective Theory, to jet physics @LHC.

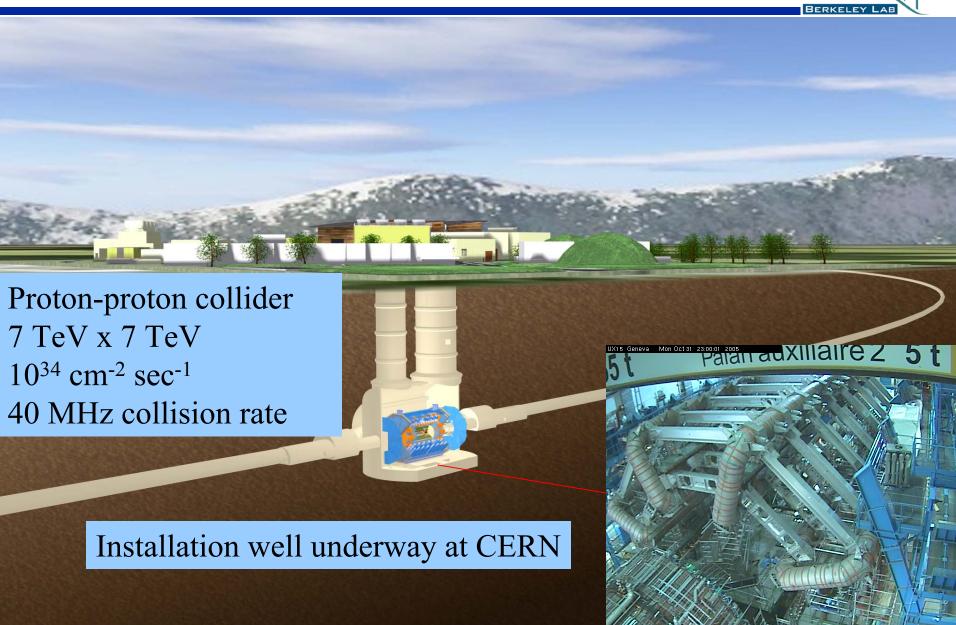
Long Term Goal: multi-loop event generator for LHC



# Future Program: The Centerpieces

## **ATLAS** at the Large Hadron Collider



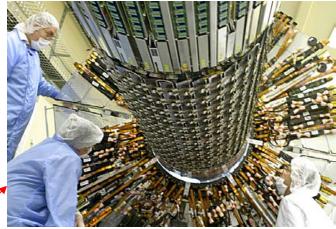


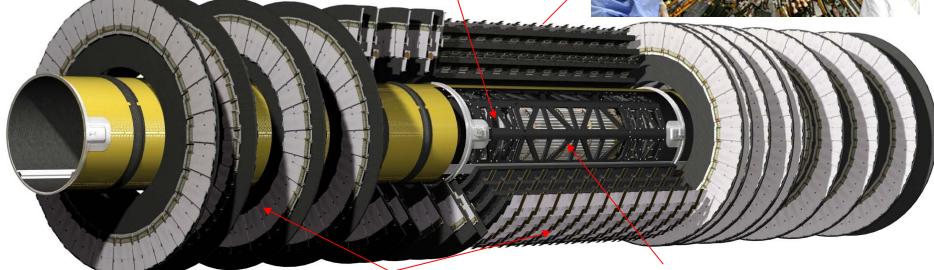
# **ATLAS Silicon Tracking Detector**



Final assembly underway







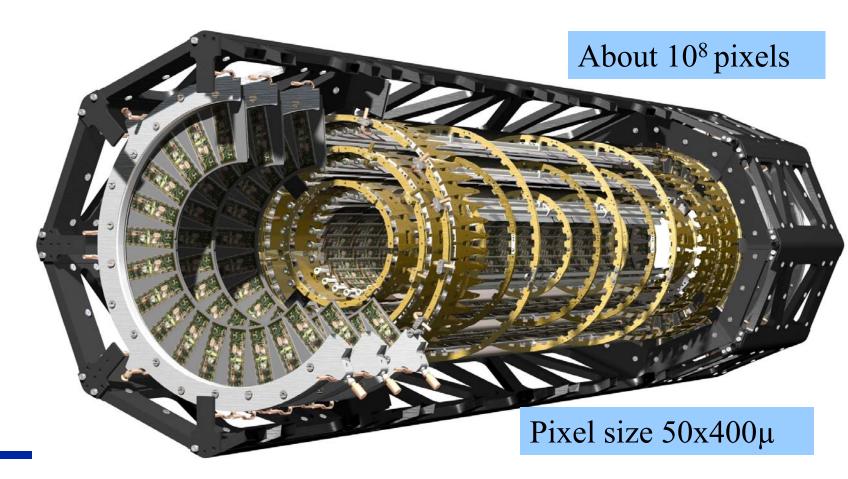
Silicon strip detector

Silicon pixel detector

#### Silicon Pixel Detector



- Completion of work at LBNL by summer 2006
- K. Einsweiler(Pixel Project Leader) and others resident at CERN



# LBNL Leadership in ATLAS Physics, Software, Simulation



- Continued leadership of ATHENA framework software.
- D. Quarrie from LBNL re-elected as Software Coordinator for ATLAS and is resident at CERN
- I. Hinchliffe continues his leadership role in Physics Coordination for ATLAS, notably in data challenges leading to the most recent ATLAS Physics Workshop this past summer in Rome.
- Recent substantial increase in LBNL work on tracking software, coordinated by M. Shapiro.
- LBNL(along with ANL and BNL) selected as Analysis Support Center for US ATLAS.

## Joint Dark Energy Mission and SNAP

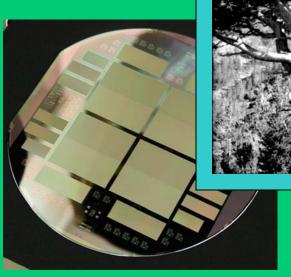


- High-level of SNAP R&D funding began in FY04
- Very substantial progress on R&D
- Expecting reduced level of funding through this fiscal year, FY06 (-7%).
- DOE continues to be very enthusiastic and supportive of JDEM and continues to push on NASA to establish a mission
- We are responding to NASA call for proposals for JDEM advanced mission concept studies due in March, 2006.
- SNAP Collaboration continues to grow
- Next SNAP collaboration meeting at FNAL in four weeks

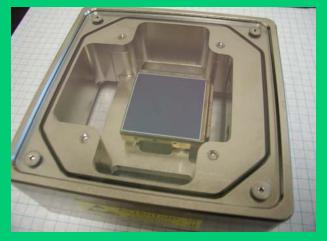
#### **Optical**

#### **CCD** electronics

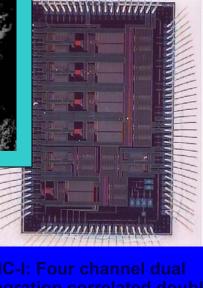
### ics



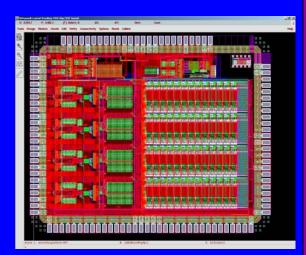
Current wafer with four SNAP CCDs – 3.5kx3.5k, 10.5 µm pixels.



Rockwell 2k x 2k HyVisi



CRIC-I: Four channel dual integration correlated double sampler, operated at 140K.

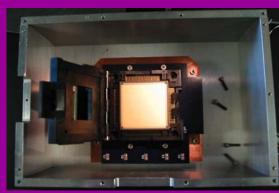


CRIC-II: with 13-b ADC

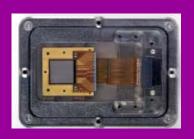
#### IR



Rockwell 2k x 2k, 1.7 um MCT.



Raytheon 1k x 1k, 1.7 um MCT



InGaAs 1k x 1k, 1.6 um looks like this.



# Future Program: In Development

# **ILC Project**



#### **Detector R&D:**

- Monolithic Pixel Sensors (LDRD)
- TPC Digital VLSI Readout
- nanoBPM: high resolution beam position monitors (LCRD)

#### **Detector Concepts:**

Leadership in LDC and GLD International detector concept studies

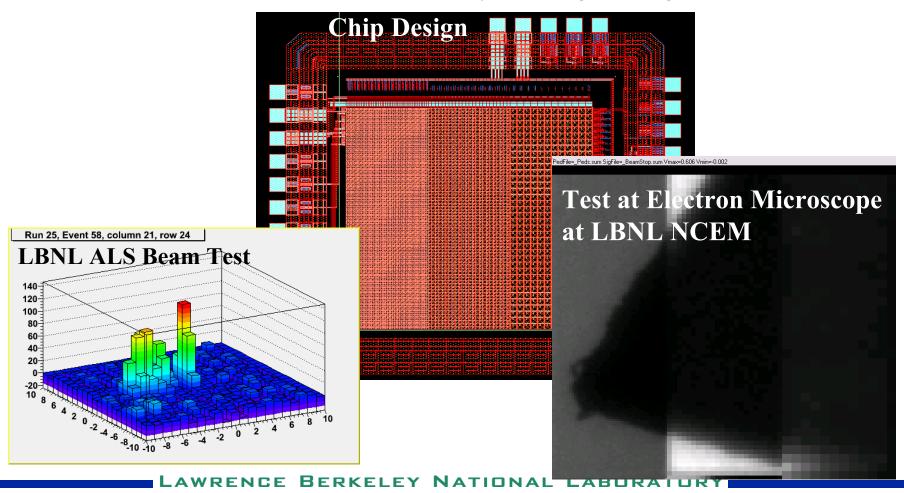
#### **Physics studies:**

- Definition of performance requirements and detector benchmarking
- Connection of ILC Physics and Cosmology (White paper on ILC and Cosmo)
- Study of interplay and complementarity with LHC physics program

# LDRD Program on Monolithic Si Pixel Sensors for ILC



Development, Characterisation and Test of new Pixel Detectors with 10 μm pixels, O(1 μm) single point resolution, on-chip data reduction and fast readout (collaborative effort of Physics, Engineering, Nuclear Sciences)



# Measurement of $\theta_{13}$ with Reactor Neutrinos



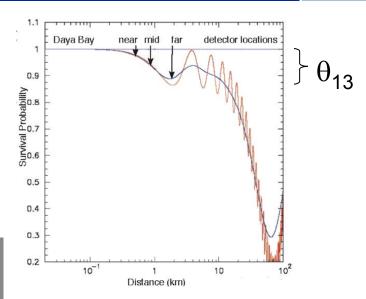
Neutrinos

$$U_{MNSP} \sim \begin{pmatrix} 0.8 & 0.5 & \mathsf{U_{e3}} \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

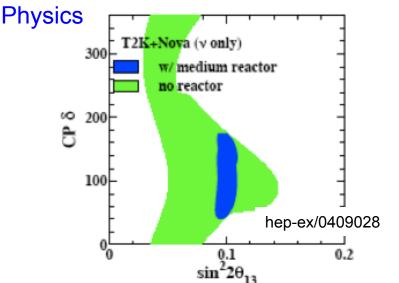
#### R&D Towards a New Reactor Experiment

- detector design
- cost estimates
- rock property testing
- PMT testing
- acrylic and scintillator testing
- Monte Carlo simulation

Daya Bay



#### Important Input to Precision Oscillation





# Community Service, Education & Outreach

# **Service to the Community**



Barnett Vice-Chair, APS Calif. Sec.; VP AAPT No. Calif. Sec.;

Chair, ATLAS Outreach

Murayama FNAL PAC; DPF Executive Committee

Roe NUSAG; FNAL PAC; URA visiting committee;

RSVP scientific assessment committee;

DESY Scientific Council; Vice Chair, DPF

Cahn HEPAP; Dark Energy Task Force;

Chair, RSVP scientific assessment committee

Siegrist MUCOG; LHC oversight

Perlmutter HEPAP; JDEM Science Definition Team

Carithers Chair, DPF

Levi JDEM Science Definition Team

Linder JDEM Science Definition Team

# Particle Data Group 50th Anniversary

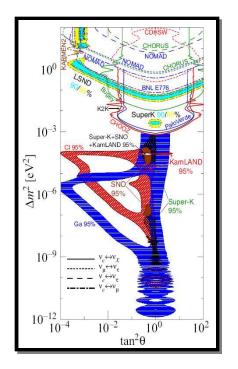




- Review of Particle Physics
- Education/Outreach Programs

Leading a Collaboration of 156 authors from 17 countries (90 institutions) + 700 contributors.





RPP: 500 new papers, 1700 new measurements, 119 reviews. 28,000 Booklets, 13,000 RPP books, website: 5-10 million hits/yr.

According to SLAC Library, RPP is the all-time top cited article in HEP with 21,500 citations (2nd is Weinberg's SM paper with 5424).



Growing coverage of Astrophysics and Cosmology

#### **Education and Outreach**



#### Involving Students, Teachers & the Public

QuarkNet sites

#### **QuarkNet – Co-Founder and Co-Pl**

Centers at 54 universities, 11 different HEP experiments, 500 high schools in 37 states. Impacts on 60,000 students/yr.

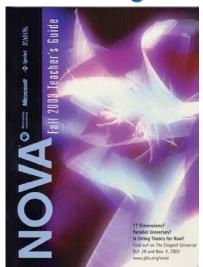
Changing teachers and teaching by making them part of of research collaborations.



#### The Particle Adventure

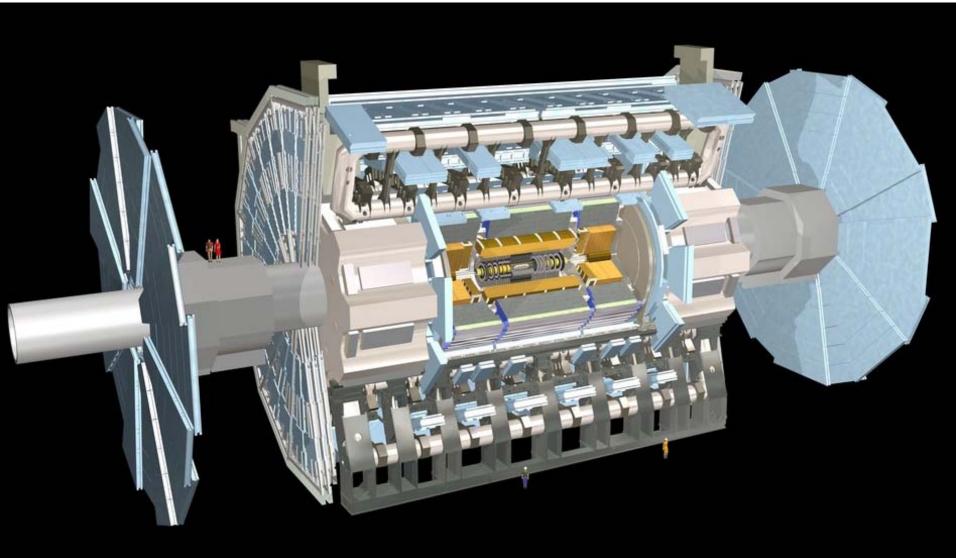
Languages: Spanish, French, Polish, Finnish, Chinese, Italian, Portuguese, Slovak, Greek, German, Norwegian, Dutch.

Featured by: Scientific American, Discovery Channel, USA Today, Education World, DOE's KidzZone, Collaboration with NOVA on "Brian Greene's Elegant Universe"

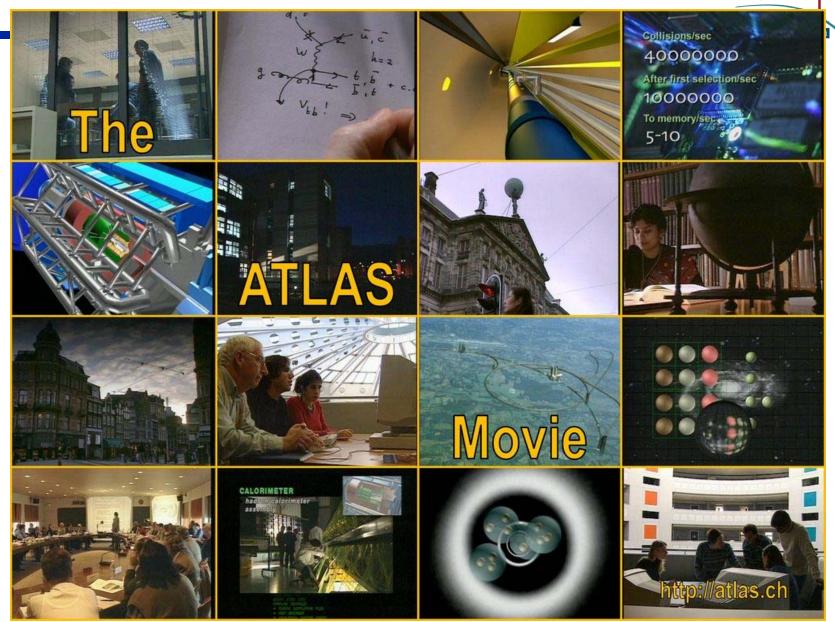


## **ATLAS Animation**





# **ATLAS Reaches Out**





# Funding and Division-wide Issues

# **Budgetary Outlook**



- FY04-FY06 budget decrease could not be absorbed without severe damage to the program
- We have reduced the size of our workforce via reduction in force during FY05
  - ~20% drop in permanent scientific staff
  - ~25% drop in administrative support costs
- Continued staff reductions in CDF and BaBar efforts
- We are reducing research efforts across the division to match funding

#### **Concerns and Risks**



# Our program is well aligned with the priorities of the field BUT

- Delays in JDEM/SNAP R&D put at risk technical readiness for project start
- Because ATLAS is under-funded, we are reaching out to the University community to seek new collaborators
- Because our supernova cosmology program is under stress, we are rethinking our long-term plans
- Because we cannot further cut ATLAS or SN
   Cosmology, we have ramped down strong analysis
   efforts in BaBar and CDF

# **Budget Information (\$K)**



	FY04	FY05	FY06	
	Actual	Allocation	President's	
LBNL Physics Research	21419	19165	18293	
ATLAS Project	3694	2434	2020	
JDEM/SNAP R&D	2498	2950	2900	
Total Funding	27611	24549	23213	

# **LDRD Support for Physics**



<u>Title</u>	<u>PI</u>	<u>2002</u>	<u>2003</u>	<u>2004</u>	<u>2005</u>	2006
						(10/05)
Foundations for a SuperNova/ Acceleration Probe (SNAP)	Levi/Perlmut	1,199,700	_	_	_	_
Modeling of High Energy Physics Detectors	Hinchliffe	94,600	99,800	_		_
POLARBEAR: An Experiment to Measure Polarization Anisotropy in the CMB	Lee	101,400	100,100			_
Future Experiments in Neutrino Physics	Freedman	X	47,000	64,000		_
Designing a Novel Reactor Neutrino Oscillation Experiment for Measuring the Unknown Mixing Angle Theta- 13		X	X	263,300	300,000	300,000
Silicon Detectors for a Linear Collider	Battaglia	x	Х	×	180,000	
New Directions for Theoretical Physics at the Tev-Scale	Murayama	х	Х	X	250,000	·

#### **Areas for Advice**



- Advise Steve Chu on the quality and impact of our work
- Advise us how to improve making our science case to the community
  - we are seeking support from our collaborations & building broader user support
  - ongoing physics at LBNL squeezed dramatically...
- How do we make the case to the lab for further investment in the Division in the post-SNAP LDRD era?
  - Neutrinos
  - Astrophysics development, theory and experiment
  - New instrumentation for future experiments and other fields
- How do we make a better enterprise with the support we have now?
- What opportunities are we missing? What are we "not seeing"?

### **Summary**



- Present program is producing great physics
- Future program is very exciting, technical progress is excellent
- New ideas are very attractive

LBNL is a center of excellence that serves the HEP community well

### **Prospects**



# We look forward to great physics!

- CP violation
- Higgs
- \*SUSY
- Dark energy
- Extra dimensions and even more